

*$^{238}\text{U}/^{235}\text{U}$  Variations in Meteorites:  
Extant  $^{247}\text{Cm}$  and Implications for  
Pb-Pb Dating*

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# Introduction

- The  $^{238}\text{U}/^{235}\text{U}$  isotope ratio has long been considered invariant in meteoritic materials (equal to 137.88).
- This assumption is a cornerstone of the high-precision Pb-Pb dates that define the absolute age of the solar system.
- In this study, the authors
  - present high-precision  $^{238}\text{U}/^{235}\text{U}$  ratios obtained from CAIs of Allende meteorite
  - demonstrate that this assumption is incorrect for these materials

## Sample: CAIs of the Allende meteorite



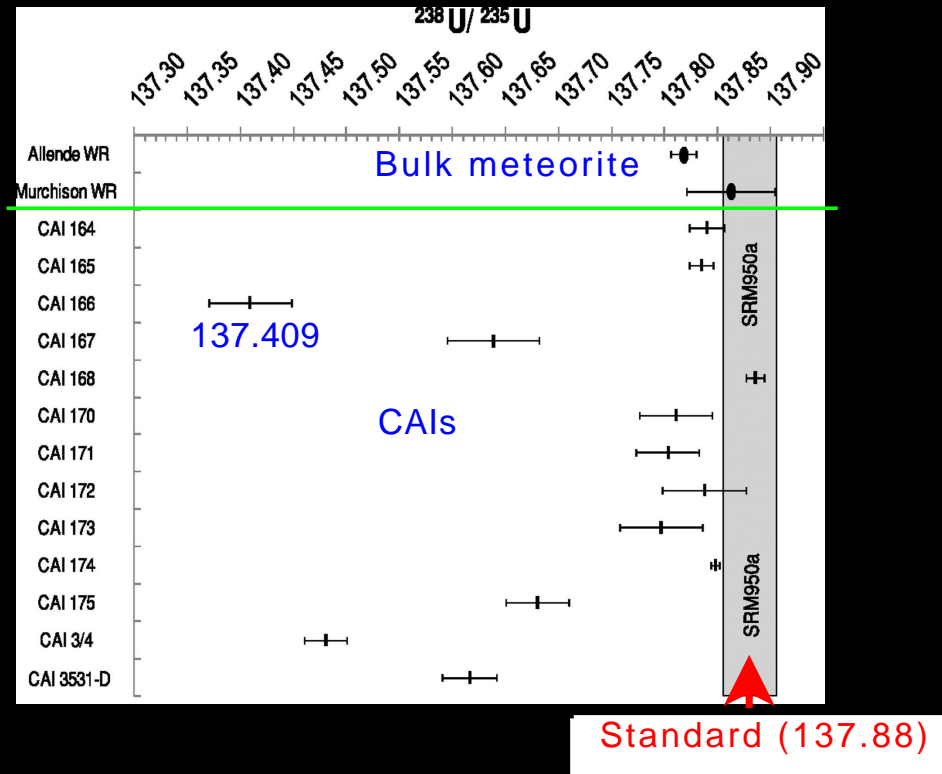
Allende meteorite - image by Matteo Chinellato

Calcium-aluminium-rich inclusions (CAIs) represent the first solids to condense from the cooling protoplanetary disk during the birth of the solar system

## Methods

- Samples (CAIs) were separated from different sections of the Allende meteorite.
- Samples were dissolved in  $\text{HNO}_3$ ,  $\text{HF}$ , and  $\text{HClO}_4$ .
- 5 % of each sample reserved for trace element measurements (REE patterns, Th/U and Nd/U ratios).
- Trace element measurements were performed with ICPMS.
- Uranium was separated with a chromatographic extraction method.
- Uranium isotope measurements were performed with MC-ICPMS.

# $^{238}\text{U}/^{235}\text{U}$ isotope values for the samples

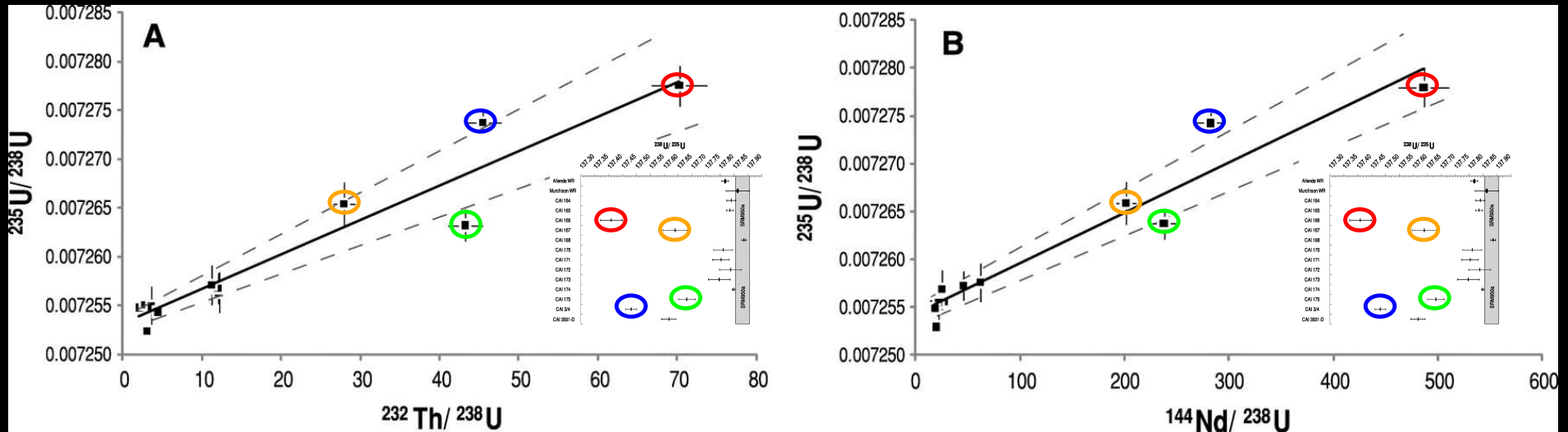


- Most CAIs differ outside uncertainties from SRM950a
- Five CAIs have significantly lower  $^{238}\text{U}/^{235}\text{U}$  values

## $^{238}\text{U}/^{235}\text{U}$ variations in the samples

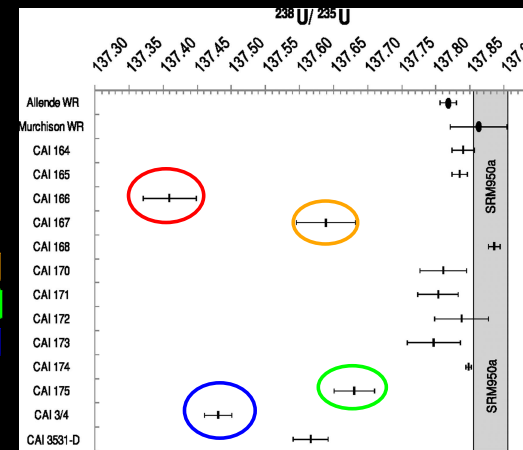
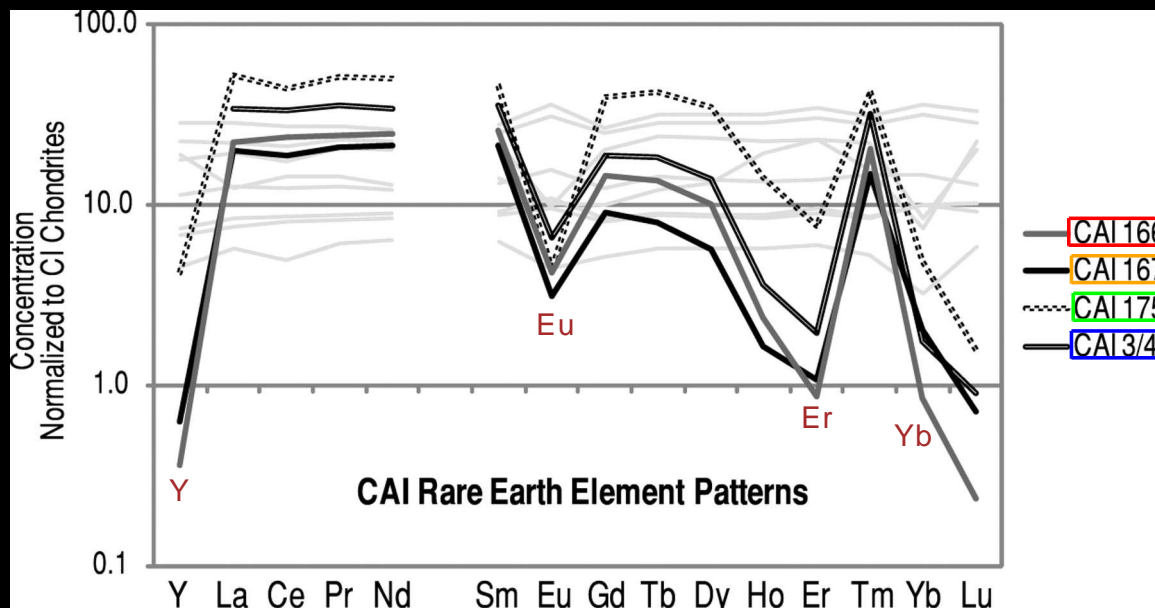
- Uranium isotope variations in meteorites may be produced by decay of extant  $^{247}\text{Cm}$  to  $^{235}\text{U}$
- $^{247}\text{Cm}$  is only created in supernova (explosion of star) and decays to  $^{235}\text{U}$  with a half-life of 15.6 My.
- If  $^{247}\text{Cm}$  decay is the primary mechanism for  $^{238}\text{U}/^{235}\text{U}$  variability, then materials with a high initial Cm/U value would contain a higher relative amount of  $^{235}\text{U}$  than those with lower initial Cm/U values.
- Because Cm has no long-lived stable isotope, the initial Cm/U ratio of a sample cannot be directly determined.
- Because Th and Nd have similar geochemical behavior to Cm, Th/U and Nd/U ratios can be used as proxies for the initial Cm/U ratio in the samples

# $^{232}\text{Th}/^{238}\text{U}$ and $^{144}\text{Nd}/^{238}\text{U}$ plotted vs $^{235}\text{U}/^{238}\text{U}$



- Large variations of the Th/U and Nd/U suggests that fractionation of Th and Nd from U occurred in the early solar nebula.
- Th/U and Nd/U ratios correlate with reciprocal values of  $^{238}\text{U}/^{235}\text{U}$
- The samples with a high initial Th/U and Nd/U values contain a higher relative amount of  $^{235}\text{U}$
- The correlation of  $^{238}\text{U}/^{235}\text{U}$  ratios with proxies for Cm/U provides strong evidence that the observed variations of  $^{238}\text{U}/^{235}\text{U}$  in CAIs were produced by the decay of extinct  $^{247}\text{Cm}$  to  $^{235}\text{U}$  in the early solar system.

# REE patterns of CAIs



- A special group of CAIs, called group II CAIs, are distinguished by a unique pattern of REE.
- The REE pattern suggests a complex condensation history involving fractional condensation.
- The four CAIs that have the highest Nd/U and Th/U ratios (as well as the lowest  $^{238}\text{U}/^{235}\text{U}$ ) are all classified as group II CAIs.



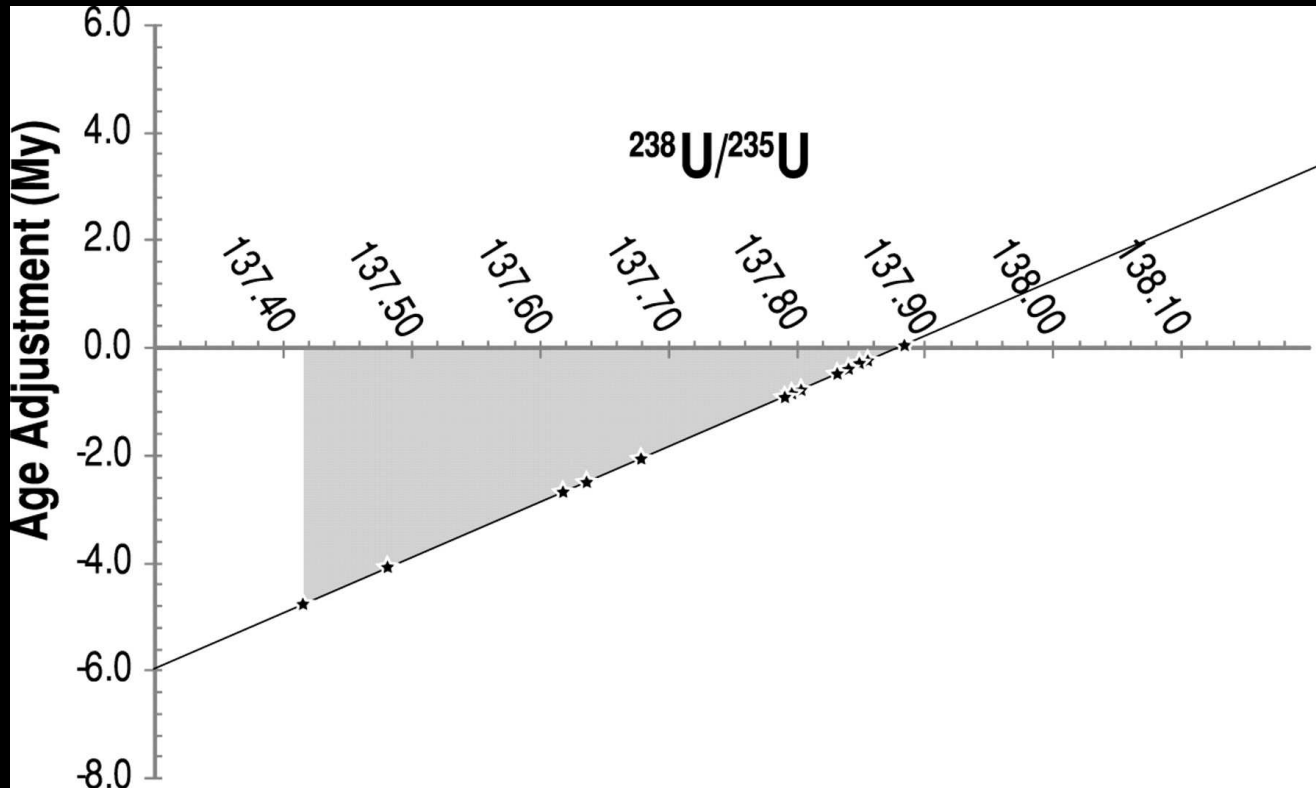
# Pb-Pb dating

- U-Pb system has half-lives and systematics suitable for resolving events in the first 10 My of the solar system.
- The Pb-Pb age equation has been used to calculate the absolute ages of both meteorites and terrestrial materials.

$$\frac{\frac{^{207}\text{Pb}}{^{204}\text{Pb}} - \left(\frac{^{207}\text{Pb}}{^{204}\text{Pb}}\right)_i}{\frac{^{206}\text{Pb}}{^{204}\text{Pb}} - \left(\frac{^{206}\text{Pb}}{^{204}\text{Pb}}\right)_i} = \left(\frac{^{235}\text{U}}{^{238}\text{U}}\right) \left(\frac{e^{\lambda_5 t} - 1}{e^{\lambda_8 t} - 1}\right) = \frac{1}{137.88} \left(\frac{e^{\lambda_5 t} - 1}{e^{\lambda_8 t} - 1}\right)$$

- $\lambda_5$ : the decay constant for  $^{235}\text{U}$  to  $^{207}\text{Pb}$  (half-lives, 0.704 By)
- $\lambda_8$ : the decay constant for  $^{238}\text{U}$  to  $^{206}\text{Pb}$  (half-lives, 4.47 By)
- All Pb-Pb ages reported today are based on the fixed  $^{238}\text{U}/^{235}\text{U}$  value.
- Any deviation of  $^{238}\text{U}/^{235}\text{U}$  from 137.88 would cause miscalculation in the determined Pb-Pb age of a sample

# Age adjustment required for samples



- $^{238}\text{U}/^{235}\text{U}$  variation results in 5My overestimation
- The largest overestimation occur in the CAls experienced the largest Cm/U fractionation

# Summary

- CAIs of the Allende meteorite display variable  $^{238}\text{U}/^{235}\text{U}$  ratios
- The correlation of  $^{238}\text{U}/^{235}\text{U}$  ratios with proxies for Cm/U provides strong evidence that the observed variations of  $^{238}\text{U}/^{235}\text{U}$  in CAIs were produced by the decay of extant  $^{247}\text{Cm}$  to  $^{235}\text{U}$  in the early solar system.
- The variable  $^{238}\text{U}/^{235}\text{U}$  ratios implies substantial uncertainties in the ages that were previously determined by Pb-Pb dating of CAIs, which may be overestimated by several million years.
- All future Pb-Pb studies must include  $^{238}\text{U}/^{235}\text{U}$  ratios.